

Fate of the universe: gauge independence and advanced precision

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Introduction



The Nobel Prize in Physics 2013 was awarded jointly to François Englert and Peter W. Higgs "for the theoretical discovery of a mechanism that contributes to our understanding of the origin of mass of subatomic particles, and which recently was confirmed through the discovery of the predicted fundamental particle, by the ATLAS and CMS experiments at CERN's Large Hadron Collider."

Higgs potential

SM Higgs sector: complex scalar doublet Φ

$$\mathcal{L} = \partial_\mu \Phi^\dagger \partial^\mu \Phi - V(|\Phi|^2), \quad V = \mu^2 |\Phi|^2 + \lambda |\Phi|^4$$

Unbroken phase: $\mu^2 > 0$

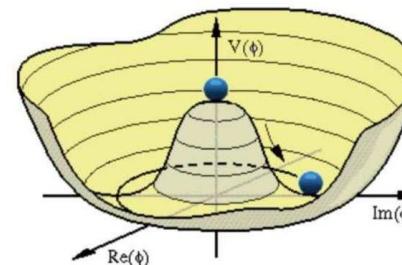
Broken phase: $\mu^2 < 0$

After SSB and Higgs mechanism:

$$\Phi = \begin{pmatrix} 0 \\ \frac{v+H}{\sqrt{2}} \end{pmatrix}$$

$$\mathcal{L} = \frac{1}{2} \partial_\mu H \partial^\mu H - V(H), \quad V = -\frac{\lambda v^4}{4} + \lambda v^2 H^2 + \lambda v H^3 + \frac{\lambda}{4} H^4$$

- $\left. \frac{\partial V}{\partial H} \right|_{H=0} = 0 \quad \rightsquigarrow \quad -\mu^2 = \lambda v^2 \equiv \frac{m_H^2}{2}$
- $\frac{gv}{2} \equiv m_W \quad \rightsquigarrow \quad v = 2^{-1/4} G_F^{-1/2} = 246.220 \text{ GeV}$
- m_H is free parameter.



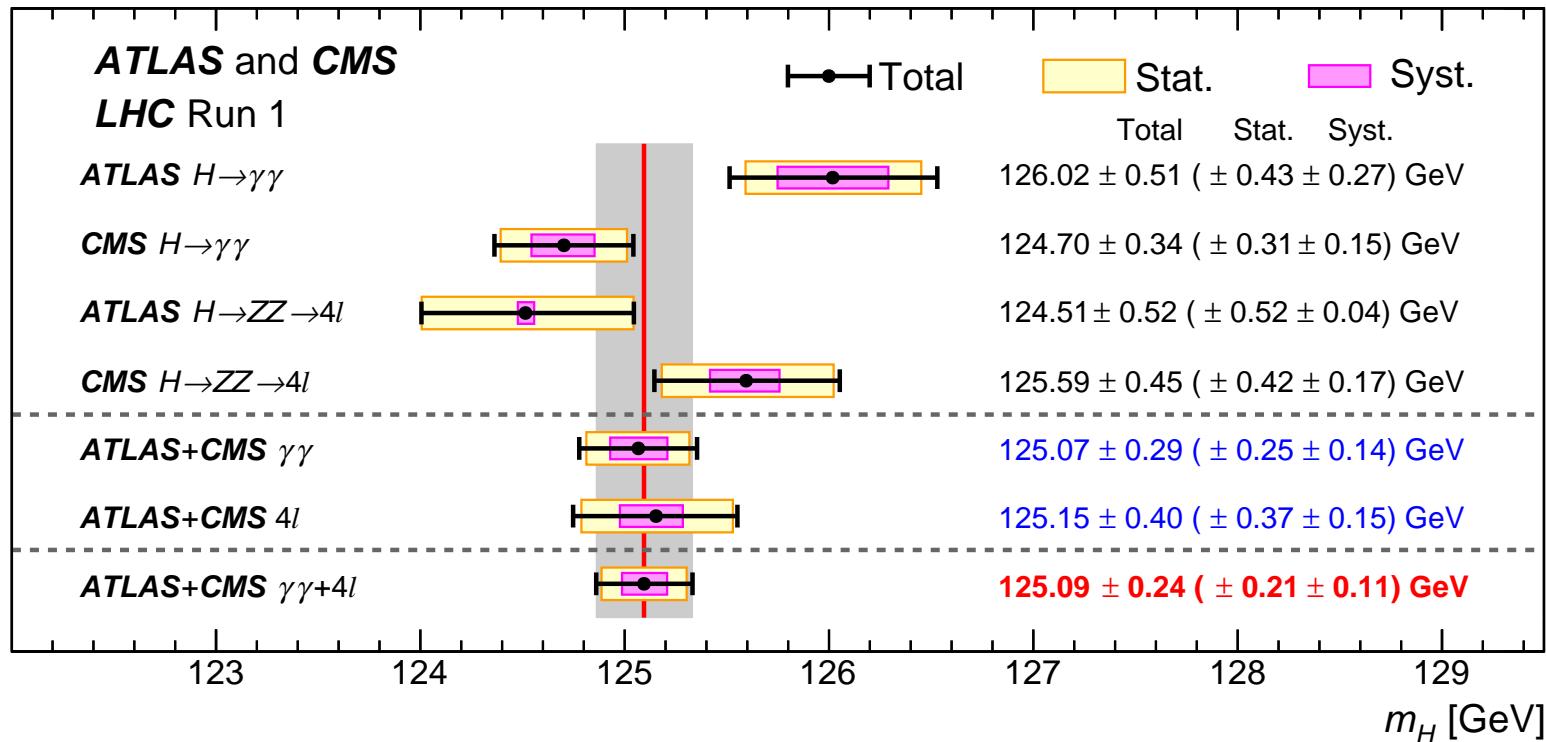
So far, **bare** fields and parameters.

Properties of the Higgs boson

$$\mathcal{L}_H = \left(m_W^2 W_\mu^+ W^{-\mu} + \frac{m_Z^2}{2} Z_\mu Z^\mu \right) \left(1 + \frac{H}{v} \right)^2 - \sum_f m_f \bar{f} f \left(1 + \frac{H}{v} \right) - \frac{m_H^2}{2} H^2 \left(1 + \frac{H}{2v} \right)^2$$

Quantum numbers	$Q = 0$ $J^{PC} = 0^{++}$
VEV	$v = 2^{-1/4} G_F^{-1/2} \approx 246.22 \text{ GeV}$
Couplings	$g_{VVH} = 2^{5/4} G_F^{1/2} m_V^2 \quad V = W, Z$ $g_{VVHH} = 2^{3/2} G_F m_V^2$ $g_{ffH} = 2^{3/4} G_F^{1/2} m_f$ $\lambda = 2^{-1/2} G_F m_H^2$ $g_{HHH} = 6v\lambda$ $g_{HHHH} = 6\lambda$
Mass	M_H is free parameter

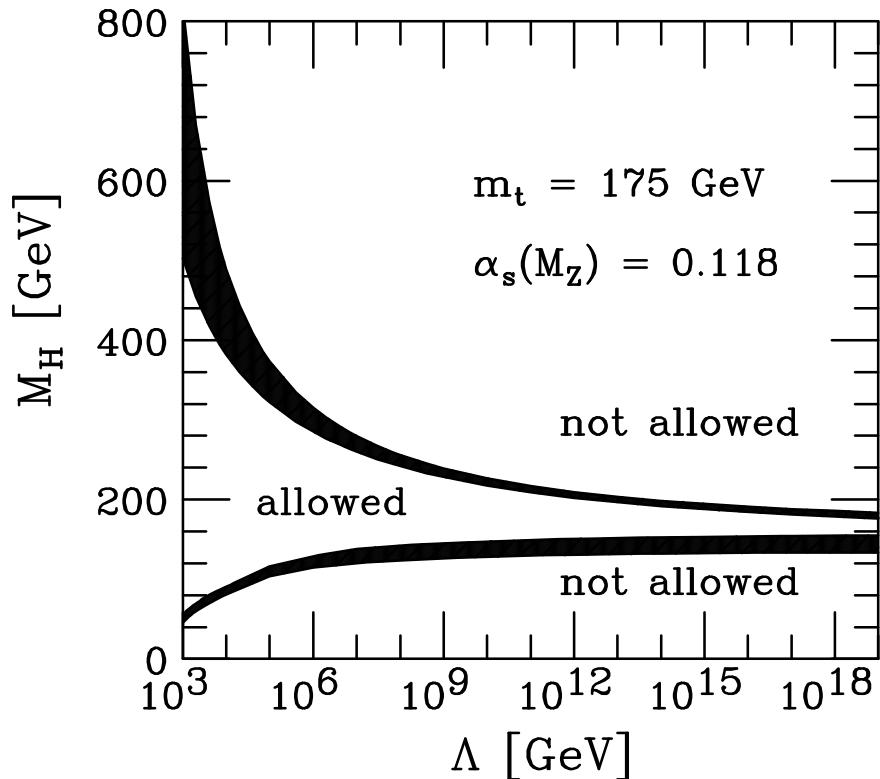
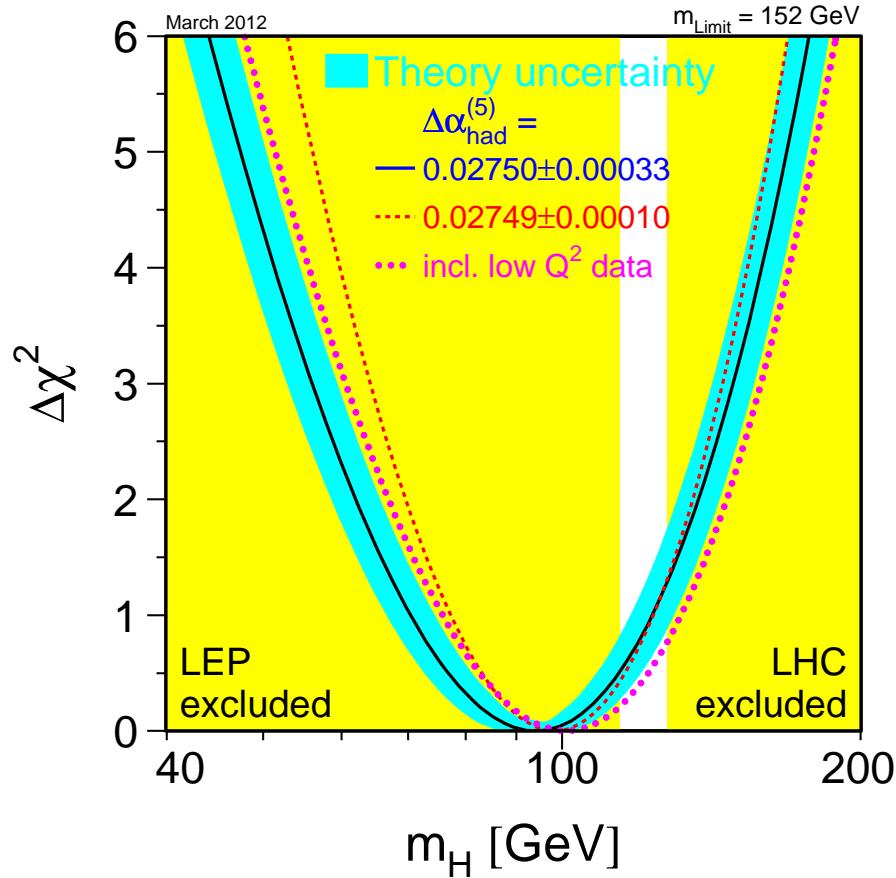
Mass measurements



$$M_H = (125.09 \pm 0.24) \text{ GeV}$$

ATLAS & CMS, PRL114(2015)191803

EW precision tests, triviality & vacuum stability



- $m_H = (125.09 \pm 0.24)$ GeV agrees w/ EW precision data.
- Triviality bound satisfied.
- How about vacuum stability bound?

Renormalization: RG evolution

Cosmological applications require reliable predictions over very large range of scales: $v \lesssim \mu \lesssim M_P$

Use $\overline{\text{MS}}$ renormalization scheme: running couplings

$\lambda(\mu), y_t(\mu), g_s(\mu), \dots$

Two-step procedure: 1. RG evolution:

$$\mu^2 \frac{d\lambda(\mu)}{d\mu^2} = \beta_\lambda = \frac{1}{16\pi^2} (12\lambda^2 + 6\lambda y_t^2 - 3y_t^4) + \dots$$

$$\mu^2 \frac{dy_t(\mu)}{d\mu^2} = \beta_{y_t} = \frac{1}{16\pi^2} y_t \left(\frac{9}{4} y_t^2 - 4g_s^2 \right) + \dots$$

$$\mu^2 \frac{dg_s(\mu)}{d\mu^2} = \beta_{g_s} = \frac{1}{16\pi^2} g_s^3 \left(-\frac{11}{2} + \frac{n_f}{3} \right) + \dots$$

$$\beta_\lambda^{(3)}, \beta_{y_t}^{(3)}$$

Chetyrkin, Zoller, JHEP06(2012)033; 04(2013)091

Bednyakov *et al.*, JHEP01(2013)017; PLB722(2013)336; NPB875(2013)552

$$\beta_{..., y_t}^{(3)}, \beta_{g_s, y_t}^{(3)}$$

Mihaila *et al.*, PRL108(2012)151602; PRD86(2012)096008

$$\beta_{g_s}^{(3)}$$

Tarasov *et al.*, PLB93(1980)429

Threshold corrections

2. Matching at $\mu = \mathcal{O}(v)$:

$$\lambda(\mu) = 2^{-1/2} G_F m_H^2 [1 + \delta_H^{(1)}(\mu) + \dots]$$

$$\delta_H^{(1)}(\mu) = \frac{G_F m_H^2}{8\pi^2 \sqrt{2}} \left[6 \ln \frac{\mu^2}{m_H^2} + \frac{25}{2} - \frac{3}{2} \pi \sqrt{3} + \mathcal{O}\left(\frac{m_Z^2}{m_H^2} \ln \frac{m_H^2}{m_Z^2}\right) \right]$$

Sirlin, Zucchini, NPB266(1986)389

$$y_t(\mu) = 2^{3/4} G_F^{1/2} m_t [1 + \delta_t^{(1)}(\mu) + \dots]$$

$$\delta_t^{(1)}(\mu) = \frac{Q_t^2 \alpha + C_F \alpha_s(\mu)}{4\pi} \left(-3 \ln \frac{\mu^2}{m_t^2} - 4 \right)$$

$$+ \frac{G_F m_t^2}{8\pi^2 \sqrt{2}} \left[\frac{9}{2} \ln \frac{\mu^2}{m_t^2} + \frac{11}{2} - 2\pi \frac{m_H}{m_t} + \mathcal{O}\left(\frac{m_H^2}{m_t^2} \ln \frac{m_t^2}{m_H^2}\right) \right]$$

Hempfling, BK, PRD51(1995)1386

$\delta_H^{(\alpha\alpha_s)}, \delta_t^{(\alpha\alpha_s)}$ Bezrukov, Kalmykov, BK, Shaposhnikov, JHEP10(2012)140

$\delta_H^{(y_t^4)}, \delta_t^{(y_t^4)}$ Degrassi *et al.*, JHEP08(2012)098; BK, Veretin, NPB885(2014)459

$\delta_H^{(\alpha^2)}, \delta_t^{(\alpha^2)}$ Buttazzo *et al.*, JHEP12(2013)089

$\delta_x^{(\alpha^2)}$ for all x BK, Veretin, Pikelner, NPB896(2015)19

$\overline{\text{MS}}$ renormalization scheme

Parameters of the symmetric phase: $g, g', \lambda, m_\phi, y_f$

Parameters of the broken phase: e, m_W, m_Z, m_H, m_f

Tree-level relationships:

$$\frac{1}{e^2} = \frac{1}{g^2} + \frac{1}{g'^2}$$

$$\frac{4m_W^2}{v^2} = g^2 \quad \frac{4m_Z^2}{v^2} = g^2 + g'^2 \quad \frac{m_H^2}{2v^2} = \lambda \quad \frac{2m_f^2}{v^2} = y_f^2$$

$$\frac{1}{v^2} = \frac{\lambda}{-m_\phi^2} = \frac{e^2}{4m_W^2(1 - m_W^2/m_Z^2)}$$

Treat as exact in the $\overline{\text{MS}}$ renormalization scheme.

On-shell renormalization scheme

- Pole masses:

$$p^2 = M_B^2 : 0 = p^2 - m_{B,0}^2 - \Pi_{BB}(p^2) \quad (B = H, W)$$

$$p^2 = M_Z^2 : 0 = p^2 - m_{Z,0}^2 - \Pi_{ZZ,T}(p^2) - \frac{\Pi_{\gamma Z,T}^2(p^2)}{p^2 - \Pi_{\gamma\gamma,T}(p^2)}$$

$$\not{p} = m_f : 0 = \not{p} - m_{f,0} - \Sigma_f(\not{p})$$

- Fine-structure constant: α_{Th} absorbs radiative corrections to Thomson scattering.

Induces large corrections $\propto \alpha \ln(q^2/m_\ell^2)$ and hadronic uncertainties! \rightsquigarrow Use instead Sirlin, PRD22(1980)971

$$G_F = \frac{\pi \alpha_{\text{Th}}}{\sqrt{2} M_W^2 (1 - M_W^2/M_Z^2)} (1 + \Delta r)$$

Matching

- Masses:

$$m_0^2 = M^2 - \Pi(M^2) = m^2(\mu) \left(1 + \frac{Z^{(1)}}{\varepsilon} + \frac{Z^{(2)}}{\varepsilon^2} + \dots \right)$$

$$Z^{(j)} = \frac{g^2}{16\pi^2} Z_\alpha^{(j)} + \frac{g^2}{16\pi^2} \frac{g_s^2}{16\pi^2} Z_{\alpha\alpha_s}^{(j)} + \left(\frac{g^2}{16\pi^2} \right)^2 Z_{\alpha^2}^{(j)} + \dots$$

- Couplings:

$$2^{1/2} G_F = \frac{1 + \Delta \bar{r}(\mu)}{v^2(\mu)}$$

$$\frac{e^2}{8m_W^2(1 - m_W^2/m_Z^2)} (1 + \Delta \bar{r}) = \left[\sqrt{Z_{2,e} Z_{2,\nu_e} Z_{2,\mu} Z_{2,\nu_\mu}} A(e + \nu_e \rightarrow \mu + \nu_\mu) \right]_{\text{hard}}$$

hard: Nullify external four-momenta and light-fermion masses before loop integration. Awramik *et al.*, PRD68(2003)053004

Threshold corrections

- Couplings:

$$g^2(\mu) = 2^{5/2} G_F M_W^2 [1 + \delta_W(\mu)]$$

$$g^2(\mu) + g'^2(\mu) = 2^{5/2} G_F M_Z^2 [1 + \delta_Z(\mu)]$$

$$e^2(\mu) = 2^{5/2} G_F M_W^2 [1 + \delta_W(\mu)] \left[1 - \frac{M_W^2}{M_Z^2} \frac{1 + \delta_W(\mu)}{1 + \delta_Z(\mu)} \right]$$

$$\lambda(\mu) = 2^{-1/2} G_F M_H^2 [1 + \delta_H(\mu)]$$

$$y_f(\mu) = 2^{3/4} G_F^{1/2} M_f [1 + \delta_f(\mu)]$$

$$g_s^2(\mu) = 4\pi\alpha_s^{(5)}(\mu) [1 + \delta_{\alpha_s}(\mu)]$$

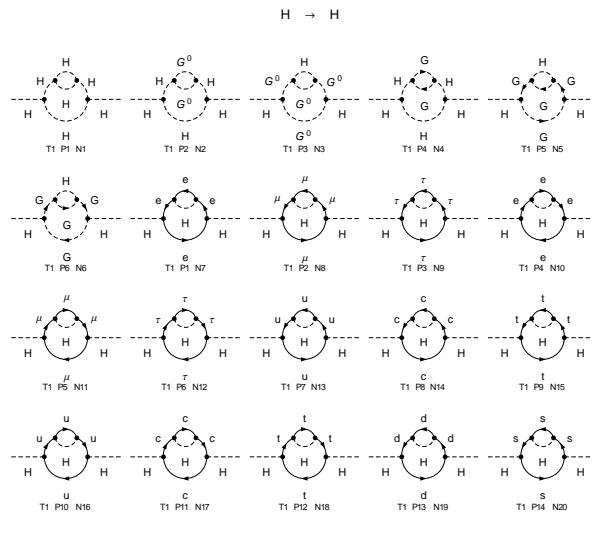
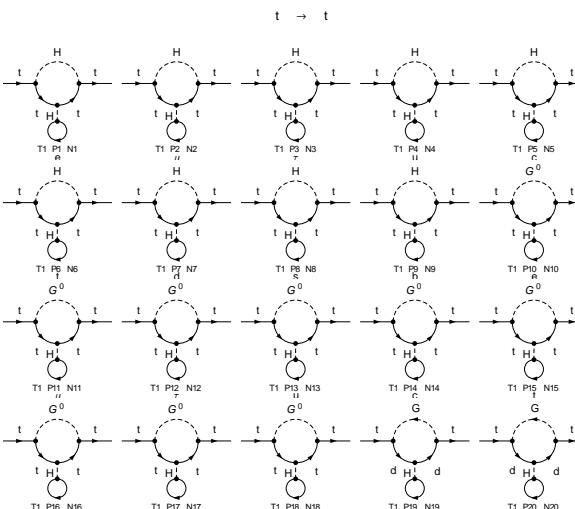
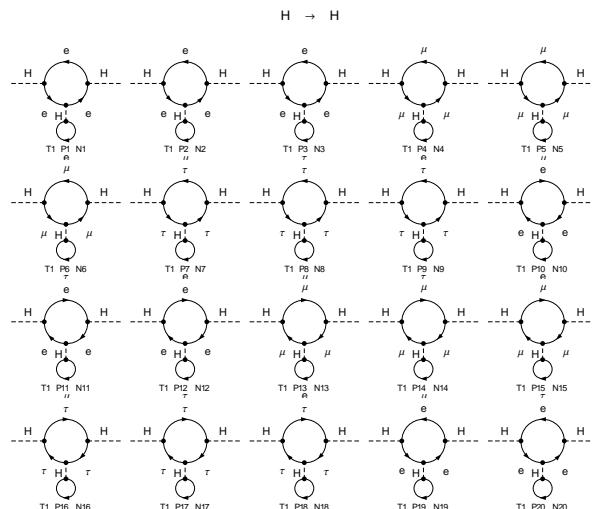
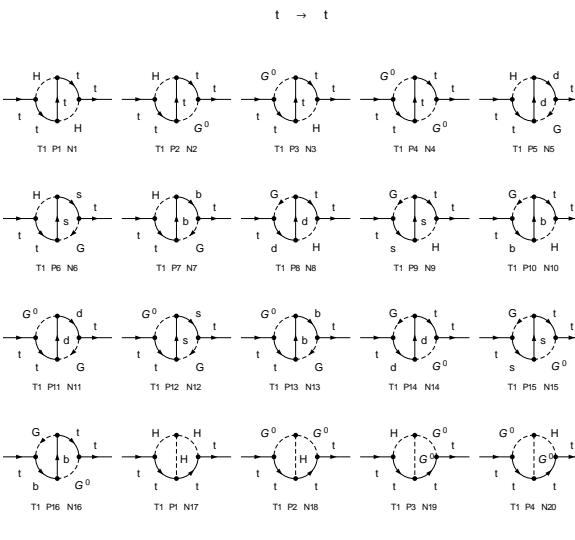
- Masses:

$$m_B^2(\mu) = M_B^2 [1 + \Delta\bar{r}(\mu)] [1 + \delta_B(\mu)] \quad B = W, Z, H$$

$$m_f(\mu) = M_f [1 + \Delta\bar{r}(\mu)]^{1/2} [1 + \delta_f(\mu)] \quad f = t, b$$

Exact two-loop results. BK, Veretin, Pikelner, NPB896(2015)19

Typical Feynman diagrams

 $\delta_H(\mu)$

 $\delta_t(\mu)$


Tools

Packages used:

- Generation of diagrams: **QGRAF**, **DIANA** Nogueira, Tentyukov
- Reduction: **TARCER** (*Mathematica*) Mertig
 - ~~> Gauge invariance upon inclusion of all tadpoles ✓
- Numerical evaluation of master integrals: **TSIL** (*C++*) Martin

Program library created: **mr** for matching and running (*C++*)

BK, Pikelner, Veretin, CPC206(2016)84

- Matching @ 2-loop EW & 4-loop QCD level
- RG evolution @ 3-loop EW & 4-loop QCD level

Available for download from URL: <http://apik.github.io/mr/>

Numerical results

- Corrections to $\delta_H(M_t)$ in 10^{-4}

M_H [GeV]	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-114.8	-107.5	-26.6 (-29.1)	-248.7
125	-114.5	-105.2	-26.4 (-29.2)	-246.1
126	-114.1	-103.1	-26.3 (-29.3)	-243.5

- Corrections to $\delta_t(M_t)$ in 10^{-4}

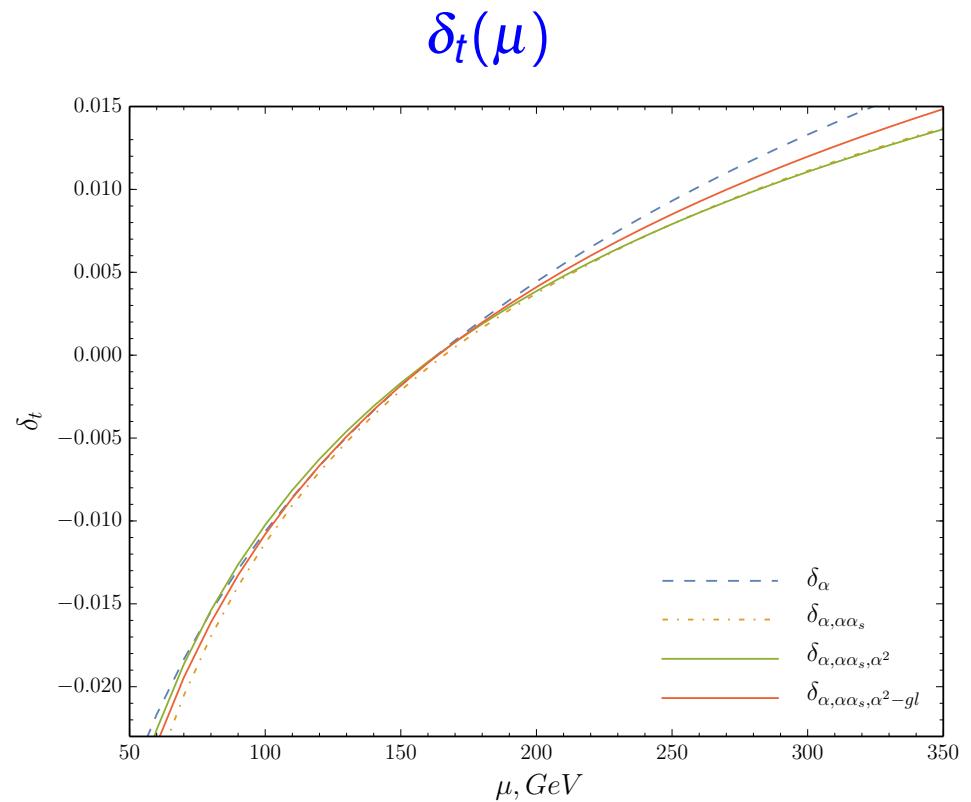
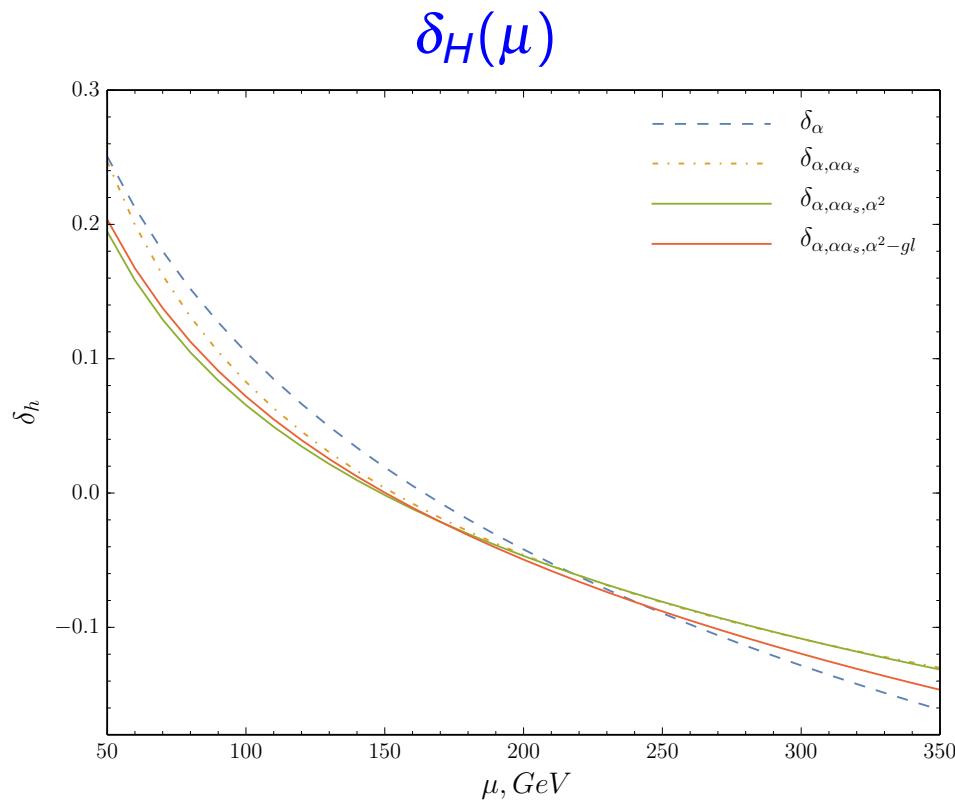
M_H [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-599.3	13.5	-4.4	2.7 (3.1)	-587.4
125	-599.3	13.2	-4.3	2.7 (3.1)	-587.7
126	-599.3	12.9	-4.2	2.7 (3.1)	-587.9

- Corrections to $\delta_b(M_b)$

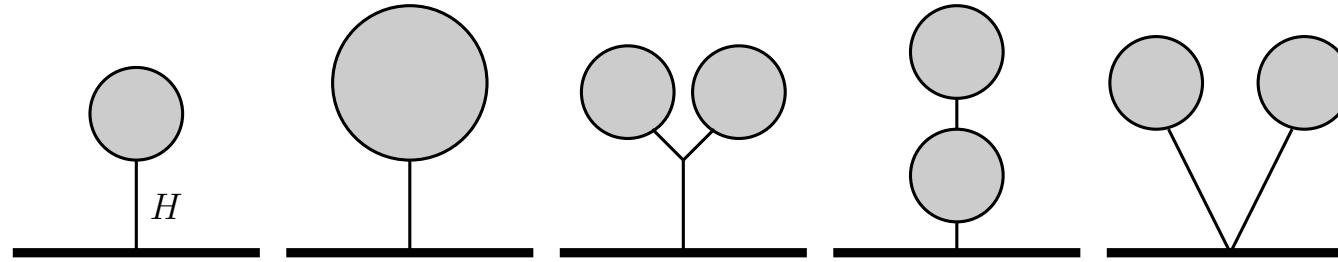
$$\begin{aligned} \{1 + \delta_b(M_b)\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ 1 - 0.1728 - 0.0190 - 0.0112 + 0.0032(0.0033) \end{aligned}$$

BK, Pikelner, Veretin, NPB896(2015)19

Numerical results



Role of tadpoles



- Tadpole is gauge dependent and $\propto 1/M_H^{2n}$ for $M_H \rightarrow 0$.
- Adjust vev $v^0 = \sqrt{-(m_\phi^0)^2/\lambda^0}$ to eliminate term $\propto H$ in bare \mathcal{L} .
Hempfling, BK, PRD51(1995)1386
- No tadpole counterterm.
- Include tadpoles order by order to ensure finiteness and gauge independence.
- $\Delta\bar{r}(\mu)$ and $\delta_x(\mu)$ are gauge independent through $\mathcal{O}(\alpha^2)$.
- At $\mathcal{O}(\alpha^2)$, $\delta_x(\mu) \propto M_H^0$ for $x = W, Z, f$; $\delta_H(\mu) \propto M_H^{-2}$; $\Delta\bar{r}(\mu) \propto M_H^{-4}$ for $M_H \rightarrow 0$.
- $m_f(\mu)$ gauge independent, but receive large EW corrections. \rightsquigarrow
Use instead *Jegerlehner, Kalmykov, BK, PLB722(2013)123*

$$m_f^Y(\mu) = 2^{-3/4} G_F^{-1/2} y_f(\mu) = M_f [1 + \delta_f(\mu)] = m_f(\mu) [1 + \Delta\bar{r}(\mu)]^{-1/2}.$$

Tadpole cancellation

- Consider $m_f(\mu)$ and $y_f(\mu)$ at $\mathcal{O}(\alpha)$ Hempfling, BK, PRD51(1995)1386

$$m_f(\mu) = M_f(1 + \delta M_f/M_f)_{\overline{\text{MS}}}$$

$$y_f(\mu) = 2^{3/4} G_F^{1/2} M_f(1 + \delta M_f/M_f - \delta v/v)_{\overline{\text{MS}}}$$

$$\delta M_f/M_f = \text{Re}[\Sigma_V^f(M_f^2) + \Sigma_S^f(M_f^2)] - 2^{1/4} G_F^{1/2} T/M_H^2$$

$$\delta v/v = [\Pi_W(0)/M_W^2 + E]/2 - 2^{1/4} G_F^{1/2} T/M_H^2$$

- Exact tadpole cancellation also in $\mathcal{O}(\alpha\alpha_s)$. Jegerlehner, Kalmykov, NPB676(2004)365; BK, Piclum, Steinhauser, NPB695(2004)199
- Incomplete tadpole cancellation in $\mathcal{O}(\alpha^2)$ BK, Veretin, NPB885(2014)459; BK, Pikelner, Veretin, NPB896(2015)19
- Similar for $\lambda(\mu)$. Sirlin, Zucchini, NPB266(1986)389; Bezrukov *et al.*, JHEP01(2012)140

Running top and bottom masses

- Corrections to $m_t(M_t) - M_t$ in GeV

M_H [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-10.38	12.08	-0.39	-0.99 (-0.47)	0.32
125	-10.38	11.88	-0.39	-0.96 (-0.45)	0.14
126	-10.38	11.67	-0.38	-0.94 (-0.44)	-0.03

- Corrections to $m_t^Y(M_t) - M_t$ in GeV

M_H [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-10.38	0.234	-0.076	0.047 (0.054)	-10.17
125	-10.38	0.229	-0.075	0.047 (0.054)	-10.18
126	-10.38	0.223	-0.073	0.047 (0.054)	-10.18

- Corrections to $m_b(M_b) - M_b$

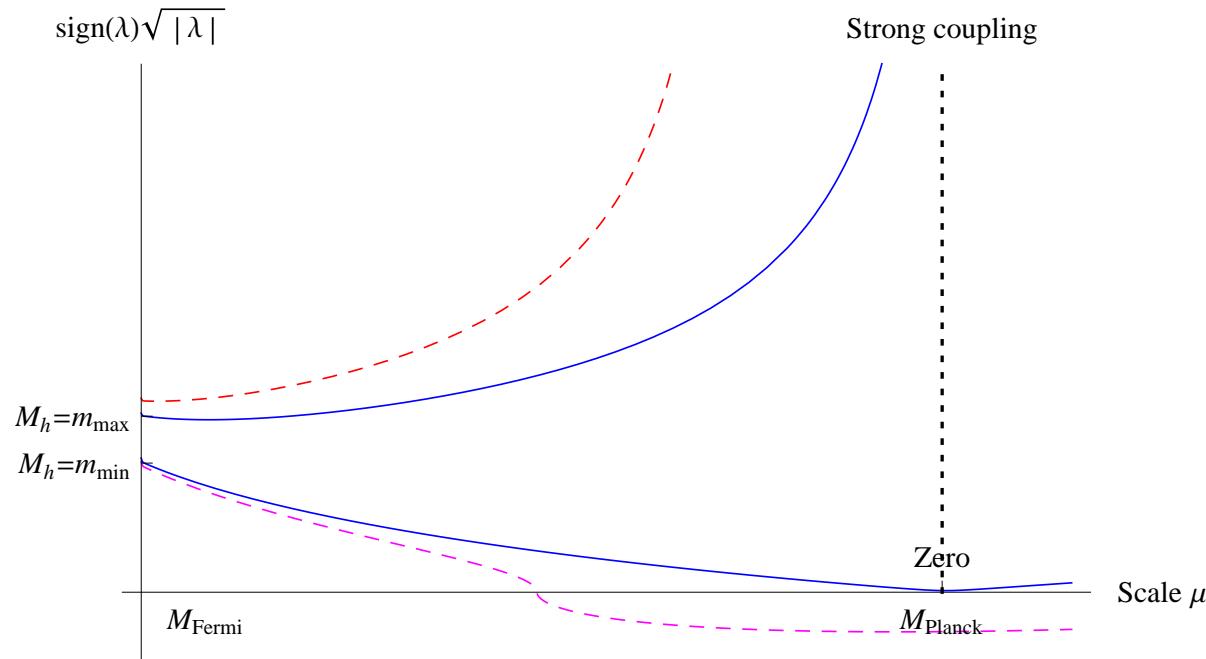
$$\begin{aligned} \{m_b(M_b) - M_b\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ -0.85 - 1.90 - 1.53 + 1.75 \quad (1.80) \text{ GeV} \end{aligned}$$

- Corrections to $m_b^Y(M_b) - M_b$

$$\begin{aligned} \{m_b^Y(M_b) - M_b\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ -0.847 - 0.093 - 0.055 + 0.016(0.016) \text{ GeV} \end{aligned}$$

↷ m_q^Y is much more perturbatively stable than $m_q(M_q)$.

Vaccum stability condition



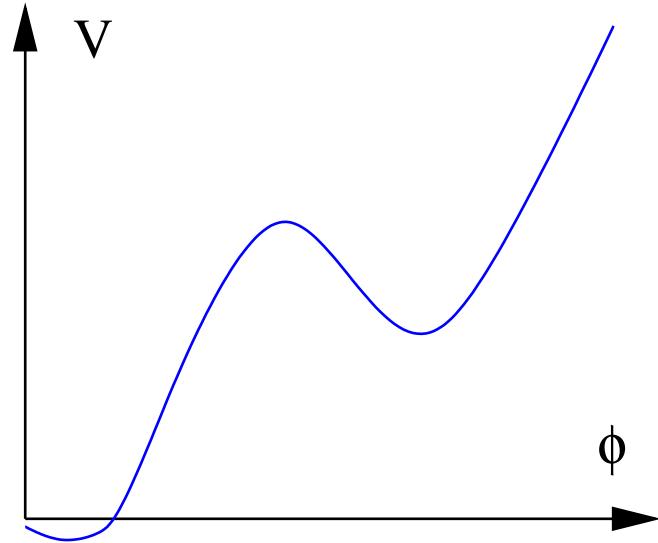
Determine μ^{cri} and M_H^{cri} for given M_t (or M_t^{cri} for given M_H) so that

$$\lambda(\mu^{\text{cri}}) = \beta_\lambda(\lambda(\mu^{\text{cri}})) = 0$$

↪ Vacuum is stable for $M_H \geq M_H^{\text{cri}}$ (or $M_t \leq M_t^{\text{cri}}$).

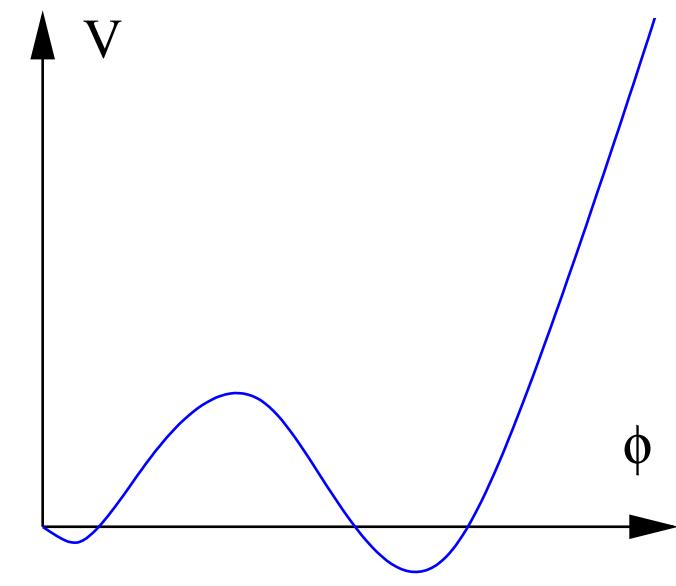
Caveat: μ^{cri} , M_H^{cri} , M_t^{cri} are gauge independent, but (slightly) scheme dependent. ↪ theoretical uncertainty

Effective potential



Fermi

Planck



Fermi

Planck

Determine $\tilde{\mu}^{\text{cri}}$ and \tilde{M}_H^{cri} for given M_t (or \tilde{M}_t^{cri} for given M_H) so that

$$V_{\text{eff}}(\tilde{\mu}^{\text{cri}}) = V_{\text{eff}}(v) \approx 0, \quad V'_{\text{eff}}(\tilde{\mu}^{\text{cri}}) = 0$$

↷ Vacuum is stable for $M_H \geq \tilde{M}_H^{\text{cri}}$ (or $M_t \leq \tilde{M}_t^{\text{cri}}$).

Caveat: $\tilde{\mu}^{\text{cri}}, \tilde{M}_H^{\text{cri}}, \tilde{M}_t^{\text{cri}}$ are gauge dependent!

Degassi *et al.*, JHEP08(2012)098; Buttazzo *et al.*, JHEP12(2013)089

Consistent approach to effective potential

- Reorganize $V_{\text{eff}}(H)$ in powers of \hbar so that expansion coefficients are gauge independent at its extrema *Andreassen et al.*, PRL113(2014)241801
- Solve $V'_{\text{eff}}(H) = 0$ for $H = \tilde{\mu}^{\text{cri}}$:

$$\lambda = \frac{1}{256\pi^2} \left[(g^2 + g'^2)^2 \left(1 - 3 \ln \frac{g^2 + g'^2}{4} \right) + 2g'^4 \left(1 - 3 \ln \frac{g'^2}{4} \right) - 48y_t^4 \left(1 - \ln \frac{y_t^2}{4} \right) \right]$$

- Require that $V_{\text{min}}^{\text{NLO}} = V_{\text{eff}}^{\text{NLO}}(\tilde{\mu}^{\text{cri}}) \geq 0$ for $M_H \geq \tilde{M}_H^{\text{cri}}$ (or $M_t \leq \tilde{M}_t^{\text{cri}}$)
e.g. in the Landau gauge
- **Caveat:** $\tilde{\mu}^{\text{cri}} > M_P$!

Critical parameters

$$X = X_0 + \Delta X_{\alpha_s} \frac{\alpha_s^{(5)}(M_Z) - \alpha_s^{(5),\text{exp}}(M_Z)}{\Delta \alpha_s^{(5),\text{exp}}(M_Z)} + \Delta X_M \frac{M - M^{\text{exp}}}{\Delta M^{\text{exp}}} \pm \delta X_{\text{par}} + \delta X_\mu^\pm \pm \delta X_{\text{tru}}$$

X	X_0	ΔX_{α_s}	ΔX_M	δX_{par}	δX_μ^+	δX_μ^-	δX_{tru}
M_t^{cri}	171.44	0.23	0.20	0.001	-0.36	0.17	-0.02
$\log_{10} \mu_t^{\text{cri}}$	17.752	-0.051	0.083	0.007	0.007	-0.006	-0.002
M_H^{cri}	129.30	-0.49	1.79	0.002	0.72	-0.33	0.04
$\log_{10} \mu_H^{\text{cri}}$	18.512	-0.158	0.381	0.008	0.173	-0.082	0.008
\tilde{M}_t^{cri}	171.64	0.23	0.20	0.001	-0.36	0.17	-0.02
$\log_{10} \tilde{\mu}_t^{\text{cri}}$	21.442	-0.059	0.094	0.005	-0.083	0.022	0.002
\tilde{M}_H^{cri}	128.90	-0.49	1.79	0.003	0.73	-0.34	0.04
$\log_{10} \tilde{\mu}_H^{\text{cri}}$	22.209	-0.181	0.436	0.007	0.092	-0.062	0.013

Importance of higher orders

- $\mathcal{O}(\alpha^2)$ corrections to all $\delta_i(\mu)$ BK, Pikelner, Veretin, NPB896(2015)19
- $\mathcal{O}(\alpha_s \alpha)$ and $\mathcal{O}(\alpha_s^4)$ corrections to $\delta_{\alpha_s}(\mu)$ Bednyakov, PLB741(2015)262; Schröder, Steinhauser, JHEP01(2006)051; Chetyrkin, Kühn, Sturm, NPB744(2006)121; BK *et al.*, PRL97(2006)042001
- $\mathcal{O}(\alpha_s^4)$ corrections to $\delta_q(\mu)$ Marquard *et al.*, PRL114(2015)142002

X	$X_0 + \delta X_\mu^\pm$	w/o $\delta_i^{O(\alpha^2)}$	w/o $\delta_{\alpha_s}^{O(\alpha \alpha_s, \alpha_s^4)}$	w/o $\delta_q^{O(\alpha_s^4)}$
M_t^{cri}	$171.44^{-0.36}_{+0.17}$	$171.55^{-0.47}_{+1.04}$	$171.43^{-0.36}_{+0.17}$	$171.24^{-0.38}_{+0.19}$
$\log_{10} \mu_t^{\text{cri}}$	$17.752^{+0.007}_{-0.006}$	$17.783^{+0.062}_{-0.008}$	$17.754^{+0.007}_{-0.006}$	$17.751^{+0.007}_{-0.007}$
M_H^{cri}	$129.30^{+0.72}_{-0.33}$	$129.06^{+0.95}_{-2.14}$	$129.32^{+0.73}_{-0.33}$	$129.72^{+0.76}_{-0.38}$
$\log_{10} \mu_H^{\text{cri}}$	$18.512^{+0.173}_{-0.082}$	$18.495^{+0.226}_{-0.531}$	$18.518^{+0.174}_{-0.082}$	$18.602^{+0.184}_{-0.094}$
\tilde{M}_t^{cri}	$171.64^{-0.36}_{+0.17}$	$171.74^{-0.46}_{+1.04}$	$171.63^{-0.36}_{+0.17}$	$171.43^{-0.37}_{+0.19}$
$\log_{10} \tilde{\mu}_t^{\text{cri}}$	$21.442^{-0.083}_{+0.022}$	$21.485^{-0.085}_{+0.343}$	$21.445^{-0.083}_{+0.022}$	$21.441^{-0.072}_{+0.014}$
\tilde{M}_H^{cri}	$128.90^{+0.73}_{-0.34}$	$128.67^{+0.95}_{-2.15}$	$128.92^{+0.73}_{-0.34}$	$129.32^{+0.76}_{-0.38}$
$\log_{10} \tilde{\mu}_H^{\text{cri}}$	$22.209^{+0.092}_{-0.062}$	$22.201^{+0.146}_{-0.171}$	$22.217^{+0.094}_{-0.062}$	$22.312^{+0.113}_{-0.082}$

Introduction
ooooo

Running & Matching
ooooooooooooooo

EW vacuum stability
oooooo●ooo

Cosmological implications
oo

Outlook
ooo

Combined results

PRL 115, 201802 (2015)

Selected for a Viewpoint in Physics
PHYSICAL REVIEW LETTERS

week ending
13 NOVEMBER 2015



Stability of the Electroweak Vacuum: Gauge Independence and Advanced Precision

A. V. Bednyakov,¹ B. A. Kniehl,² A. F. Pikelner,² and O. L. Veretin²

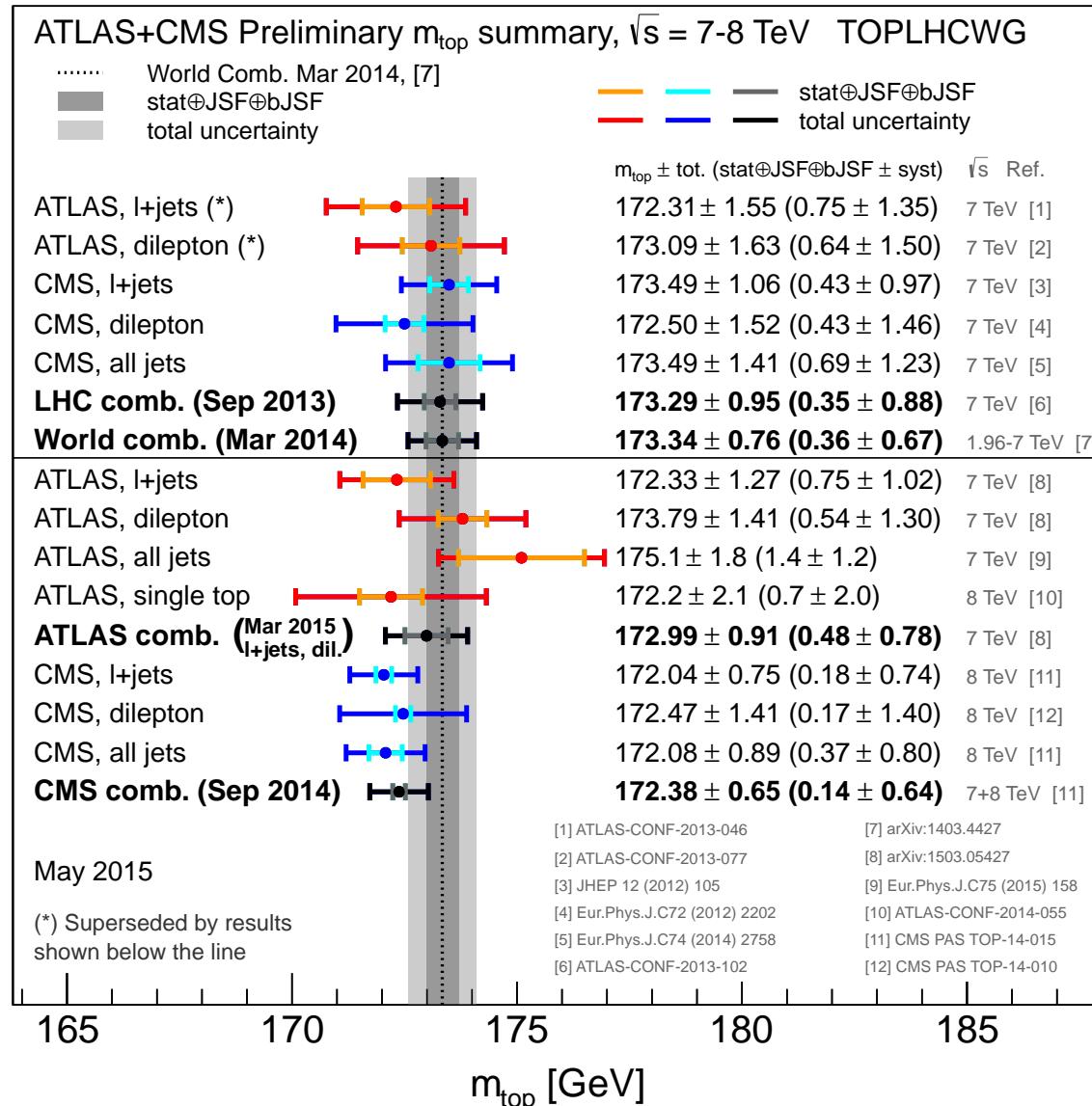
¹Joint Institute for Nuclear Research, 141980 Dubna, Russia

²II. Institut für Theoretische Physik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Germany

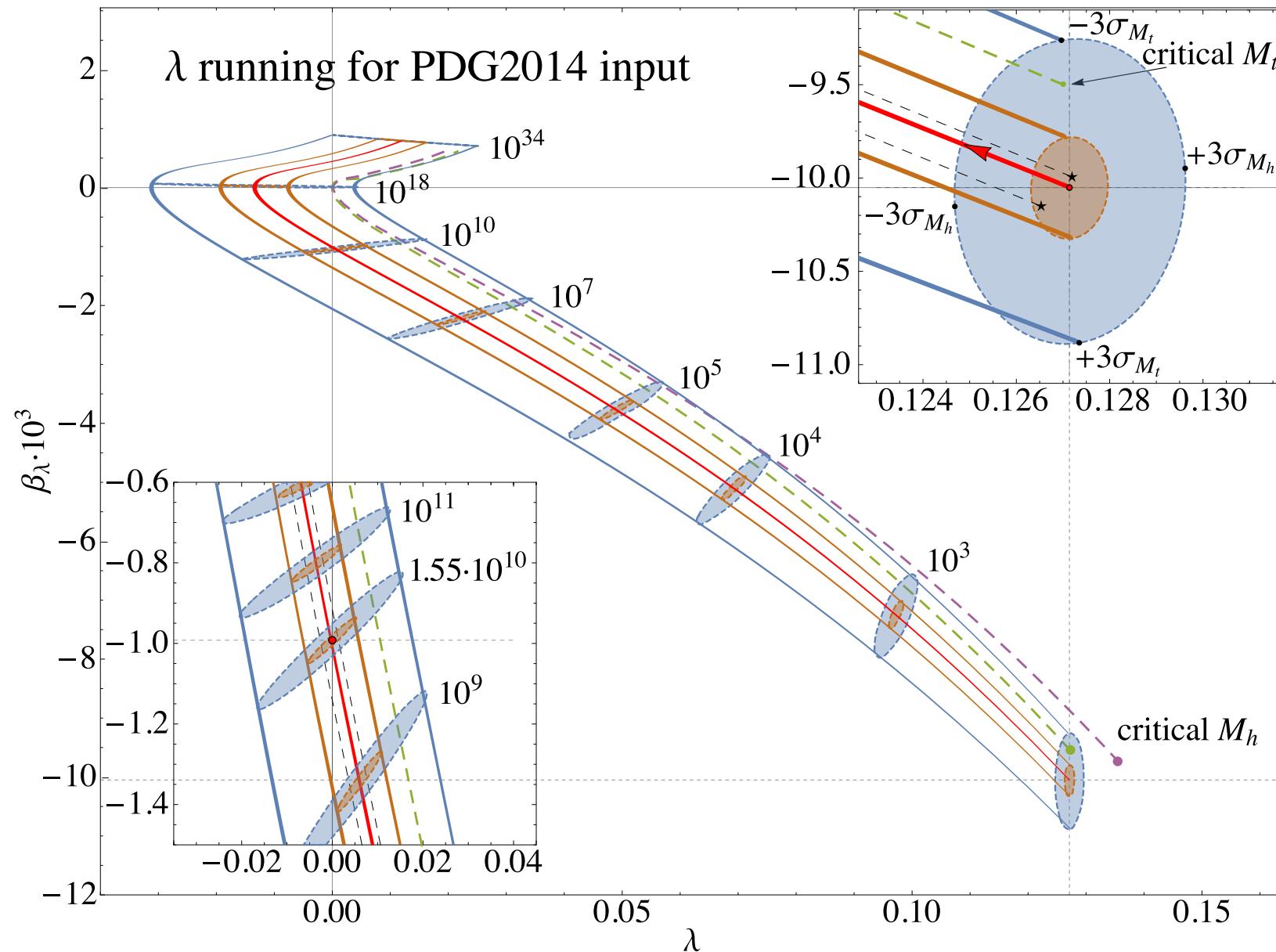
(Received 30 July 2015; revised manuscript received 24 August 2015; published 9 November 2015)

- From $\lambda(\mu)$: $M_t^{\text{cri}} = (171.44 \pm 0.30)^{+0.17}_{-0.36}$ GeV
- From $V_{\text{eff}}(H)$: $\tilde{M}_t^{\text{cri}} = (171.64 \pm 0.30)^{+0.17}_{-0.36}$ GeV
- Combination: $\hat{M}_t^{\text{cri}} = (171.54 \pm 0.30)^{+0.26}_{-0.41}$ GeV
- Experiment: $M_t^{\text{MC}} = (172.38 \pm 0.66)$ GeV ATLAS & CMS,
[arXiv:1512.02244 \[hep-ex\]](https://arxiv.org/abs/1512.02244)

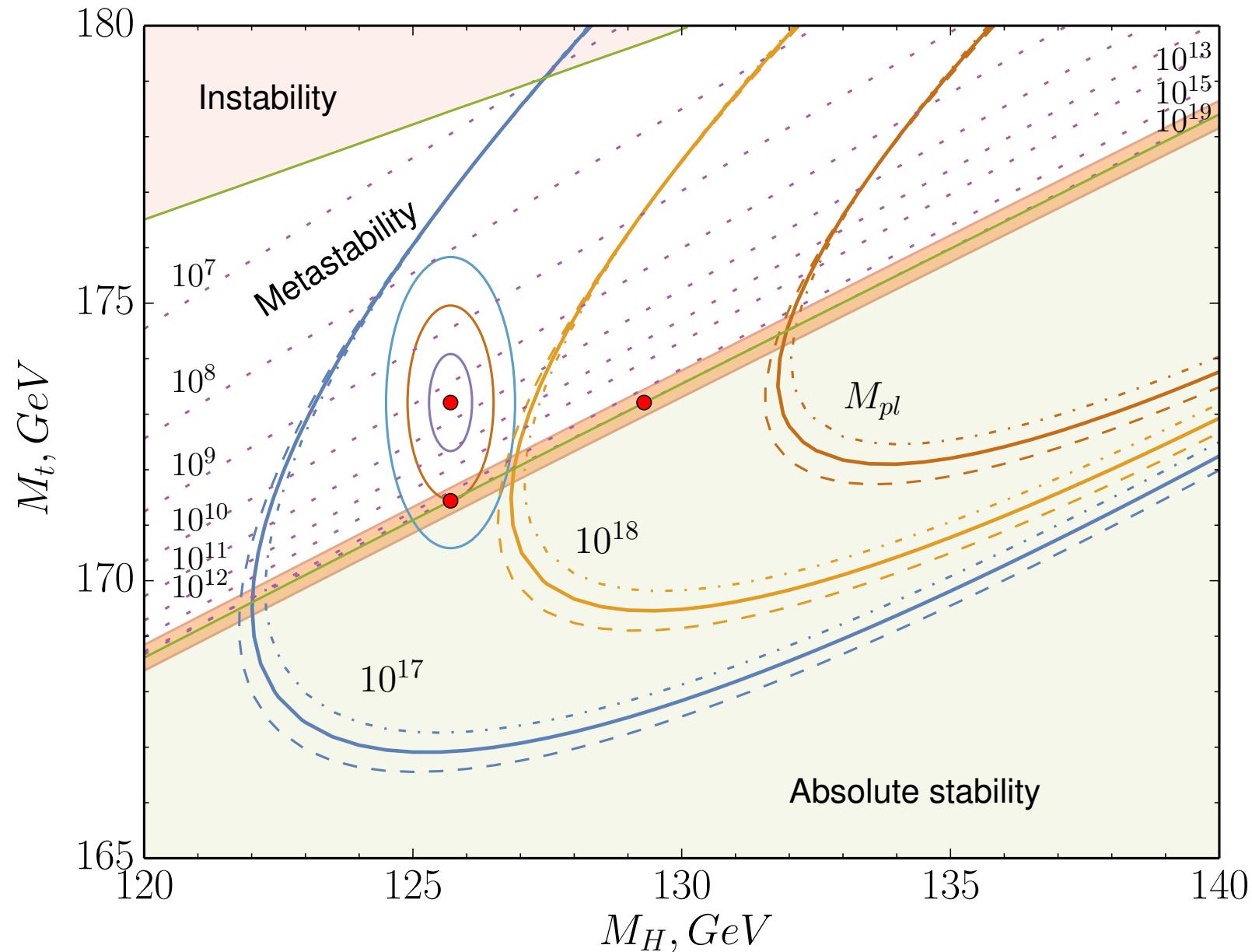
M_t Measurements



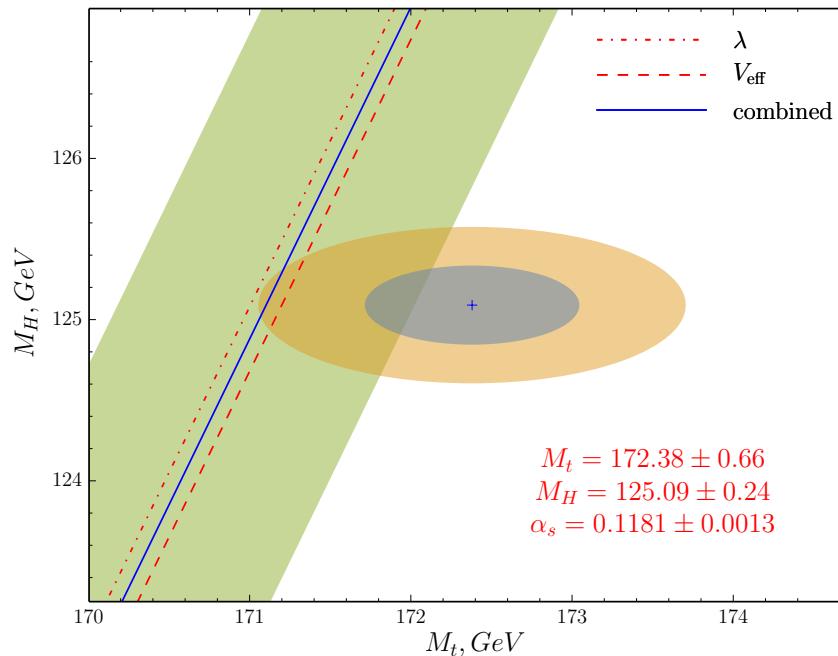
RG flow



Phase diagram



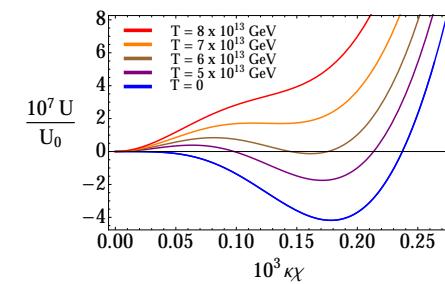
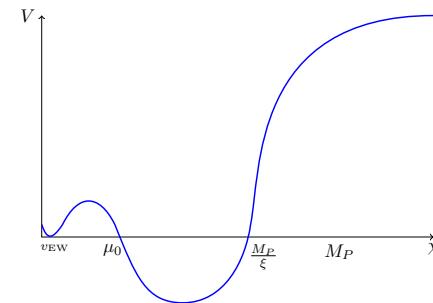
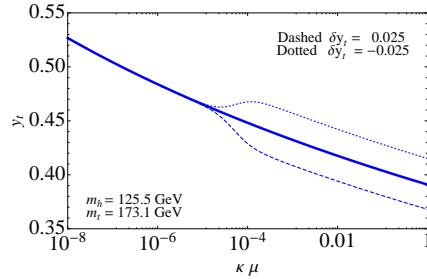
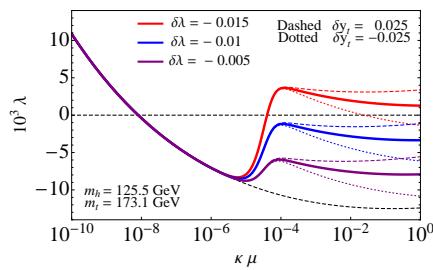
SM stable all the way up to M_P ?



- Intriguing conspiracy of SM particle masses ↗
 $\mu^{cri} \approx M_P = 1.22 \times 10^{18} \text{ GeV}$
- μ^{cri} stable w.r.t. parametric and higher-order uncertainties due to **asymptotic safety**
- Relationship between M_P and SM parameters?
- Electroweak scale determined by Planck scale physics?
- Implicit reduction of fundamental couplings?

Higgs inflaton

- Higgs field, nonminimally coupled to gravity with strength ξ , can be responsible for inflation
- Successful scenario possible even if EW vacuum is metastable
- Effective renormalization of SM couplings at scale M_P/ξ
- Symmetry restoration after inflation due to high- T effects temporarily eliminating vacuum at $H \approx M_P$

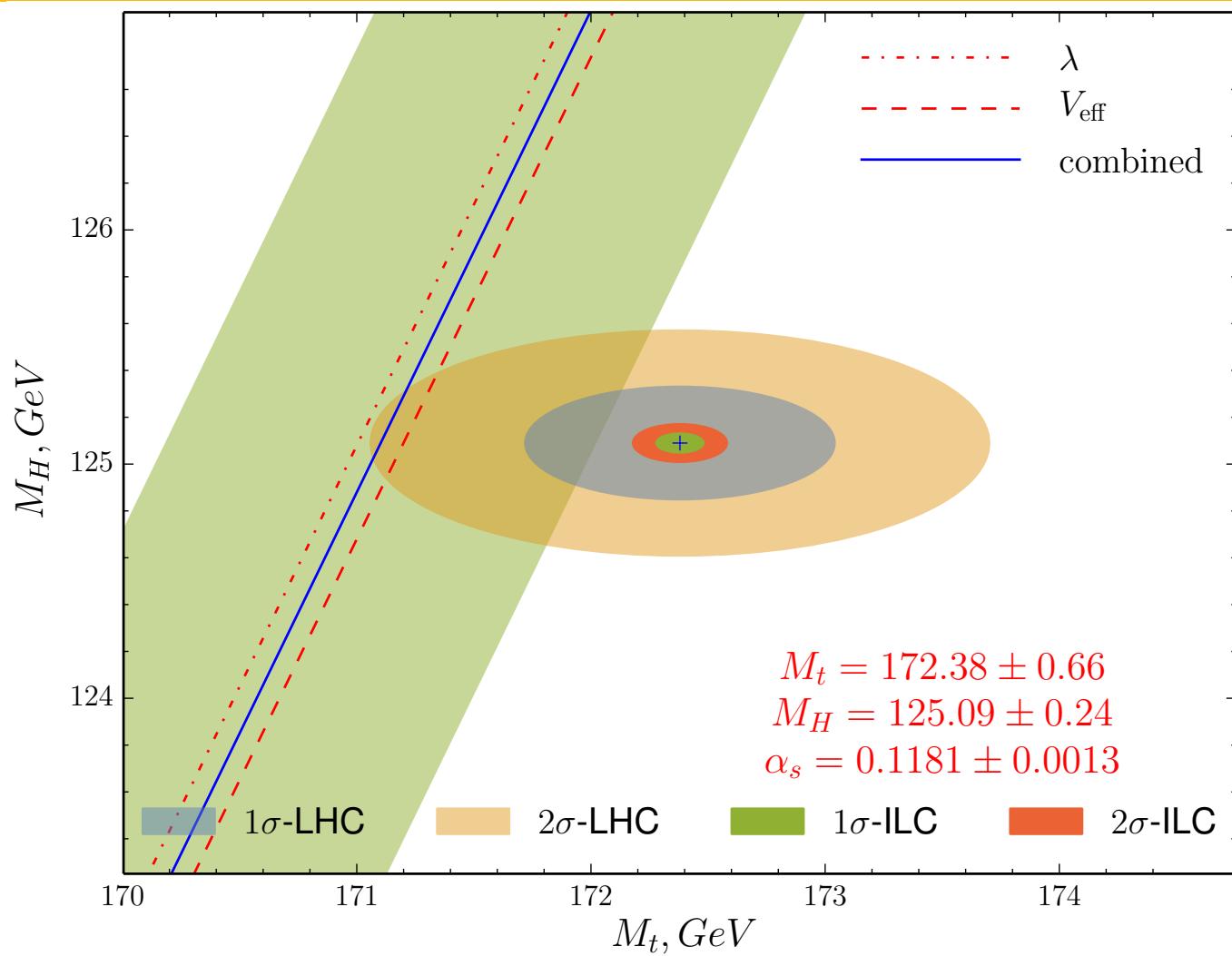


Bezrukov *et al.*, PRD92(2015)083512

Outlook: pole mass M_t

- PDG value $M_X(t \rightarrow X) = (173.21 \pm 0.87)$ GeV is **not** pole mass M_t , but just parameter in MC programs w/o RC to partonic cross sections.
- Rigorous determination of $\overline{\text{MS}}$ mass $m_t(\mu)$ from $\sigma_{\text{tot}}(p\bar{p}, pp \rightarrow t\bar{t} + X)$:
 $M_t = (170.4 \pm 1.2)$ GeV ABMP16, arXiv:1701.05838 [hep-ph]

ILC as top and Higgs factory



Anticipated errors $\delta M_t = 100$ MeV, $\delta M_H = 40$ MeV Moortgat-Pick *et al.*,
EPJC75(2015)371

BSM physics

- Depending on future precision measurements of M_H , M_t , α_s and higher-loop RC calculations, SM may be stable all the way up to M_P .
- BSM physics still necessary to solve open problems, e.g.
 - smallness of neutrino masses
 - strong CP problem
 - dark matter
 - baryon asymmetry of universe
 - unification with gravity
- Higgs portals?